

Spin ordering at weak coupling : V. M. STOJANOVIĆ
spin density waves

Antiferromagnetism has the remarkable feature of being realizable both in the strong-coupling and in the weak-coupling regimes (needless to say, by 'coupling' we here mean e-e interaction). For comparison, superconductors appear only in the weak-coupling, while crystals are realized only in the strong-coupling regime.

In both cases the starting point is the Hubbard model

$$H = -t \sum_{i,\delta} (C_{i\delta}^\dagger C_{i+\delta} + h.c.) + U \sum_i n_{i\uparrow} n_{i\downarrow}.$$

In our treatment of the strong-coupling regime we used a perturbation theory with (t/U) being the small parameter. Now we want to address the regime where kinetic energy dominates over the interactions (i.e., the weak-coupling regime). We will address this regime using what is called Hartree-Fock mean-field theory.

The existence of Néel order, mathematically expressed as

$$\langle S_i^z \rangle = \frac{1}{2} (\langle n_{i\uparrow} \rangle - \langle n_{i\downarrow} \rangle) \neq 0,$$

is anticipated within this mean-field approach.

We introduce operators $\delta O_{i\beta} \equiv n_{i\beta} - \underbrace{\langle n_{i\beta} \rangle}_{O_{i\beta}}$

$$\Rightarrow \boxed{n_{i\beta} = O_{i\beta} + \delta O_{i\beta}}$$

Using the Hartree-type decoupling of the interaction term, and to remind you (we already used it in the context of the RPA approximation) this is the decoupling of the type

$$\begin{array}{ccc}
 A & B & \longrightarrow \langle A \rangle B + A \langle B \rangle - \langle A \rangle \langle B \rangle, \\
 \swarrow & \searrow & \\
 a_{\mu}^{\dagger} a_{\mu} & b_{\nu}^{\dagger} b_{\nu} &
 \end{array}$$

we obtain the mean-field Hamiltonian

$$H^{MF} = -t \sum_{i, \delta, \beta} (C_{i, \delta}^{\dagger} C_{i+\delta, \beta} + \text{h.c.}) \\
 + U \sum_i (O_{i\uparrow} n_{i\downarrow} + O_{i\downarrow} n_{i\uparrow} - O_{i\uparrow} O_{i\downarrow})$$

the e-e-interaction effects are now absorbed in simple single-electron scattering terms

By introducing
$$\begin{aligned}
 \Omega_i^z &= \frac{1}{2} (O_{i\uparrow} - O_{i\downarrow}) \\
 \bar{n}_i &= \frac{1}{2} (O_{i\uparrow} + O_{i\downarrow}),
 \end{aligned}$$

we have
$$\begin{aligned}
 O_{i\uparrow} &= \bar{n}_i + \Omega_i^z, \\
 O_{i\downarrow} &= \bar{n}_i - \Omega_i^z,
 \end{aligned}$$

which, when substituted in the last mean-field Hamiltonian leads to

$$H_{MF} = -t \sum_{i, \delta, \beta} (C_{i\beta}^\dagger C_{i+\delta, \beta} + h.c.)$$

$$+ U \sum_i \left[\bar{n}_i (n_{i\uparrow} + n_{i\downarrow}) - \Omega_i^z (n_{i\uparrow} - n_{i\downarrow}) - \bar{n}_i^2 + (\Omega_i^z)^2 \right]$$

Is something perhaps wrong with the last mean-field Hamiltonian?

When doing mean-field theory we would ideally want our mean-field Hamiltonian to obey the same symmetries as the original Hubbard Hamiltonian.

One of these symmetries is the spin-rotational symmetry which we discussed before. H'_{MF} breaks this symmetry!

When we are decoupling a quartic (interaction) term with all four operators of the same type (i.e., describing interactions between the same kind of particles) then another decoupling (known as Fock-type) has to be taken into account:

$$U \sum_i n_{i\uparrow} n_{i\downarrow} \rightarrow -U \sum_i \underbrace{C_{i\uparrow}^\dagger C_{i\downarrow}}_{S_i^+} \underbrace{C_{i\downarrow}^\dagger C_{i\uparrow}}_{S_i^-}$$

$$= -U \sum_i (\Omega_i^+ + \delta \Omega_i^+) (\Omega_i^- + \delta \Omega_i^-)$$

All in all, the interaction term in the Hartree-Fock mean-field theory becomes

$$U \sum_i n_{i\uparrow} n_{i\downarrow} \mapsto U \sum_i \left[\bar{n}_i (n_{i\uparrow} + n_{i\downarrow}) \right.$$

$$- \Omega_i^z (n_{i\uparrow} - n_{i\downarrow})$$

$$- \Omega_i^+ c_{i\downarrow}^+ c_{i\uparrow}$$

$$- \Omega_i^- c_{i\uparrow}^+ c_{i\downarrow}$$

$$- \bar{n}_i^2 + (\Omega_i^z)^2$$

$$+ \Omega_i^+ \Omega_i^- + \text{fluct.} \left. \right]$$

Having arrived at this Hartree-Fock mean-field Hamiltonian, the further procedure entails the following:

1) Choose a set of classical variables

\bar{n}_i and Ω_i (they can in principle be different on each lattice site)

2) Each set of these variables defines an independent-electron problem, which we want to solve.

3) Find the minimum classical energy in the classical configuration space. This is the sum of the N_e (= total electron number) lowest single-particle energies (found in step 2), plus the potential energy

$$U (-\bar{n}_i^2 + \Omega_i^2)$$

We illustrate this for the half-filled Hubbard model on a cubic lattice.

Assuming that the ground state corresponds to the Néel state with a uniform charge density, we adopt

$$\bar{n}_i = \bar{n} \quad (\forall i),$$

$$\Omega_i = (0, 0, \Omega) \quad (i \in A),$$

$$\Omega_i = (0, 0, -\Omega) \quad (i \in B),$$

where \bar{n} and Ω are still to be determined.

From our mean-field Hamiltonian we can see that spin- \uparrow electrons experience potential $U^*(\bar{n} - \Omega)$ on the A sublattice and $U^*(\bar{n} + \Omega)$ on the B sublattice. For spin- \downarrow electrons, the role of the A and B sublattices is reversed:

$$H_{\text{Néel}}^{\text{MF}} = -t \sum_{i, \delta} (C_{i\delta}^\dagger C_{i+\delta} + \text{h.c.})$$

$$+ U \sum_{i \in A, B} \left\{ \bar{n} (n_{i\uparrow} + n_{i\downarrow}) \mp \Omega (n_{i\uparrow} - n_{i\downarrow}) - \bar{n}^2 + \Omega^2 \right\}$$

sign here depends on the sublattice!

The last Hamiltonian defines a "band-structure-like" problem on a lattice with two inequivalent sites per unit cell.

To solve such a problem, we first perform a Fourier transformation to momentum space. We obtain the Hamiltonian

$$H_{\text{Néel}}^{\text{MF}} = \sum_{\mathbf{k}} \left[2t \cos\left(\frac{k}{2}\right) \sum_{\delta} (C_{A, \mathbf{k}\delta}^{\dagger} C_{B, \mathbf{k}\delta} + \text{h.c.}) \right. \\ \left. + U(\bar{n} - \Omega) (C_{A, \mathbf{k}\uparrow}^{\dagger} C_{A, \mathbf{k}\uparrow} + C_{B, \mathbf{k}\downarrow}^{\dagger} C_{B, \mathbf{k}\downarrow}) \right. \\ \left. + U(\bar{n} + \Omega) (C_{A, \mathbf{k}\downarrow}^{\dagger} C_{A, \mathbf{k}\downarrow} + C_{B, \mathbf{k}\uparrow}^{\dagger} C_{B, \mathbf{k}\uparrow}) \right. \\ \left. + U(\Omega^2 - \bar{n}^2) \right],$$

where we introduced "A" and "B" fermions (that is, fermions living on two different sublattices).

Through a fermion Bogolubov transformation

$$\gamma_{\mathbf{k}\uparrow}^{\dagger} = \cos \phi_{\mathbf{k}} C_{A, \mathbf{k}\uparrow}^{\dagger} + \sin \phi_{\mathbf{k}} C_{B, \mathbf{k}\uparrow}^{\dagger},$$

$$\bar{\gamma}_{\mathbf{k}\uparrow}^{\dagger} = \sin \phi_{\mathbf{k}} C_{A, \mathbf{k}\uparrow}^{\dagger} - \cos \phi_{\mathbf{k}} C_{B, \mathbf{k}\uparrow}^{\dagger},$$

$$\gamma_{\mathbf{k}\downarrow}^{\dagger} = \cos \phi_{\mathbf{k}} C_{B, \mathbf{k}\downarrow}^{\dagger} + \sin \phi_{\mathbf{k}} C_{A, \mathbf{k}\downarrow}^{\dagger},$$

$$\bar{\gamma}_{\mathbf{k}\downarrow}^{\dagger} = \sin \phi_{\mathbf{k}} C_{B, \mathbf{k}\downarrow}^{\dagger} - \cos \phi_{\mathbf{k}} C_{A, \mathbf{k}\downarrow}^{\dagger},$$

where $\gamma_{\mathbf{k}\delta}^{\dagger}$ are the low-lying bonding states (lower-band states) and $\bar{\gamma}_{\mathbf{k}\delta}^{\dagger}$ higher-lying unoccupied states, after diagonalization we obtain

$$H_{\text{Néel}}^{\text{MF}} = \sum_{\mathbf{k}\delta} \left[\omega_{\mathbf{k}}^{-}(\bar{n}, \Omega) \gamma_{\mathbf{k}\delta}^{\dagger} \gamma_{\mathbf{k}\delta} + \omega_{\mathbf{k}}^{+}(\bar{n}, \Omega) \bar{\gamma}_{\mathbf{k}\delta}^{\dagger} \bar{\gamma}_{\mathbf{k}\delta} \right] \\ + NU(\Omega^2 - \bar{n}^2)$$

with $\omega_{\mathbf{k}}^{\pm} = U\bar{n} \pm \sqrt{U^2\Omega^2 + 4t^2 \cos^2 \frac{\mathbf{k}}{2}}$. (**)

For large $\frac{U\Omega}{t}$ the lower band corresponds to the up- (down-) electrons being predominantly on the A- (B) sublattice.

In accordance with the "prescription" given above the classical energy is given by

$$E(\bar{n}, \Omega) = \frac{\langle H_{MF} \rangle}{N} = \frac{1}{N} \sum_{\mathbf{k}=-\pi}^{\pi} \omega_{\mathbf{k}}^{-}(\bar{n}, \Omega) + U(\Omega^2 - \bar{n}^2).$$

Minimization with respect to \bar{n} and Ω leads to the equations

$$\frac{\partial E}{\partial \bar{n}} = -2\bar{n}U + \frac{1}{N} \sum_{\mathbf{k}} \frac{\partial}{\partial \bar{n}} \omega_{\mathbf{k}}^{-}(\bar{n}, \Omega) = 0$$

$$\frac{\partial E}{\partial \Omega} = 2\Omega U + \frac{1}{N} \sum_{\mathbf{k}} \frac{\partial}{\partial \Omega} \omega_{\mathbf{k}}^{-}(\bar{n}, \Omega) = 0$$

Combining this last equation with Eq. (**) above we get an implicit equation for the order parameter amplitude Ω :

$$\frac{1}{2N} \sum_{\mathbf{k}} \frac{1}{\sqrt{U^2\Omega^2 + 4t^2 \cos^2 \frac{\mathbf{k}}{2}}} = \frac{1}{U},$$

or, equivalently,

$$(\Delta) \quad \frac{1}{2\pi} \int_0^{\pi} \frac{dk}{\sqrt{1 + \left(\frac{2t}{U\Omega}\right)^2 \cos^2 \frac{k}{2}}} = \Omega.$$

Such an equation can be solved by iteration.

Let us consider the limit $\frac{U}{t} \rightarrow 0$ in one dimension. It can be expected that only the electron states close to E_F are affected; accordingly, their dispersion can be linearized

$$\omega_k = 2t \cos\left(\frac{k}{2}\right) \rightarrow \omega_k = v_F (k - k_F)$$

k_F here is

$$k_F = \frac{\pi}{2a}$$

Fermi momentum of the folded zone

A high energy cut-off momentum k_0' is chosen in such a way as to reproduce the bandwidth
($k' = k - k_F$)

$$v_F k_0' = 2t$$

With these simplifying assumptions Eq. (Δ) above reduces to

$$\frac{1}{U} = \frac{1}{2\pi v_F} \int_0^{k_0'} \frac{dk'}{\sqrt{k'^2 + \frac{U^2 \Omega^2}{4t^2}}} = \frac{1}{2\pi v_F} \sinh^{-1} \left(\frac{U v_F k_0'}{2t} \right),$$

which is equivalent to

$$U\Omega = \Delta = \frac{|v_F k_0|}{\sinh\left(\frac{2J|U_F|}{U}\right)} \approx 4t e^{-\frac{2Jt}{U}}$$

$U\Omega$ has the meaning of ^a half the gap that opens at the boundary of the folded Brillouin zone. This gap is finite for any $U > 0$.

Despite being small, the gap is present even in the weak-coupling limit. As long as the gap spans the Fermi surface (NESTING, a situation that is characteristic of 1d systems, and rarely takes place in $d > 1$) it is possible to think of Wannier states of localized spins; their difference from localized spins in strong coupling is their much larger spatial extent (coherence length) ξ , which can be estimated as

$$\frac{\xi}{a} \sim \frac{W}{2U\Omega}.$$

Because the gap $2U\Omega$ is a very small fraction of the bandwidth W , the coherence length is quite large, e.g., $\xi \sim 1000a$. This resulting weak-coupling antiferromagnet is called a spin-density wave (SDW). — it is something like a Fermi fluid with a

a wave of spin-density superimposed on it.

The SDW's occur in 1d not only at half-filling but also at other filling fractions; the nesting then happens at different wave numbers than $\frac{\pi}{2a}$ and instead of a doubling of a unit cell the SDW can ^{even} be found with wavelengths incommensurate with the original lattice period.

Since in higher dimensions ($d > 1$) nesting occurs only in very special cases (e.g., a Hubbard model on a square lattice with nearest-neighbor hopping only) the SDW's are special to the one-dimensional systems.

In nature, they are typically realized as quasi-1d systems. There is, for instance, a large class of quasi-1d organic systems which are metallic at room temperature and undergo a transition to SDW at some lower temperature. An example is furnished by the so-called Bechgaard salts.